## 802.

## Problem 53.20 (RHK)

A neutron star can be analyzed by techniques similar to those used for ordinary metals. In this case the neutrons (rather than electrons) obey the Fermi-Dirac probability function. We have to consider a neutron star of 2.00 solar masses with a radius of 10.0 km and calculate the Fermi energy of the neutrons.

## **Solution:**

Mass of a neutron,  $m_n = 1.67 \times 10^{-27}$  kg,

Mass of a neutron star of 2.00 solar masses,  $M = 2 \times 1.99 \times 10^{30}$  kg,

Radius of the neutron star,  $R = 10.0 \times 10^3$  m. Number of neutrons in the neutron star,

$$N = \frac{2 \times 1.99 \times 10^{30}}{1.67 \times 10^{-27}} = 2.383 \times 10^{57}.$$

From the above data we note that the number of neutron per cubic meter in the neutron star,

$$n = \frac{N}{\left(4\pi R^3/3\right)} = \frac{2.383 \times 10^{57}}{\left(4\pi \left(10^4\right)^3/3\right)} \text{ neutrons per m}^3$$
$$= 5.688 \times 10^{44} \text{ neutrons per m}^3.$$

At absolute zero Fermi energy is the highest energy of the occupied state. It is given by the relation

$$E_F = \frac{h^2}{8m} \left(\frac{3n}{\pi}\right)^{2/3}$$

Therefore, the Fermi energy of the neutrons in the neutron star of 2 solar masses and radius of 10 km will be

$$E_F = \frac{\left(6.63 \times 10^{-34}\right)^2}{8 \times 1.67 \times 10^{-27}} \times \left(\frac{3 \times 5.688 \times 10^{44}}{\pi}\right)^{\frac{2}{3}} \text{ J}$$
  
= 21.9×10<sup>-12</sup> J  
= 21.9×10<sup>-12</sup> × 6.242×10<sup>12</sup> MeV  
= 136.7 MeV.