

92.

Problem 16.30 (RHK)

It is conjectured that a “burned-out” star could collapse to a “gravitational radius”, defined as the radius for which the work needed to remove an object of mass m from the star’s surface to infinity equals the rest energy mc^2 of the object. We have to show that the gravitational radius of the Sun is GM_s/c^2 and determine its value in terms of the Sun’s present radius.



Solution:

Let us consider an object of mass M and surface radius R . Work required to remove an object of mass m from the surface of the object of mass M to infinity is the negative of the potential energy of the m - M system. It is

$$W_{\text{escape}} = \frac{GMm}{R}.$$

Gravitational radius of an object of mass M is the radius R_g which is such as would require energy mc^2 for an object of mass m to escape to infinity. That is

$$\frac{GMm}{R_g} = mc^2,$$

or

$$R_g = \frac{GM}{c^2}.$$

We calculate next the *gravitational radius* of the Sun.

$$M_s = 1.99 \times 10^{30} \text{ kg},$$

$$\text{Gravitationa constant } G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}.$$

We use this data and find that the *gravitational radius* of the Sun is

$$R_g (\text{Sun}) = \frac{6.67 \times 10^{-11} \times 1.99 \times 10^{30}}{9 \times 10^{16}} \text{ m} = 1.47 \text{ km}.$$

According to the Einstein's theory of gravitation the correct value of R_g is $2GM/c^2$ and is called the Schwarzschild radius.